

The structure of the maximal development for shock-forming 3D compressible Euler solutions

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with Leo Abbrescia

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3D compressible Euler flow

$$\partial_t \varrho + \partial_a(\varrho v^a) = 0,$$

$$\varrho \mathbf{B} v^i = -\partial_i p \quad (= \partial_t(\varrho v^i) + \partial_a(\varrho v^a v^i))$$

$$\mathbf{B} \mathbf{s} = 0$$

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- Two propagation phenomena: sound waves and transporting of vorticity/entropy
- Neither phenomena nor their coupling are apparent

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- Set up the **shock development problem**

Remarks on 1D theory

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- Challis (1848)
- Stokes (1850s)
- Riemann (1860)
- Oleinik (1959)
- Zabusky (1962)
- Lax (1964)
- Glimm (1965)
- Keller–Ting (1966)
- Dafermos (1970)
- Smoller (1970)
- Liu (1974)
- John (1974)
- Klainerman–Majda (1980)
- Jenssen (2000)
- Chen–Feldman (2003)
- Bianchini–Bressan (2005)
- ...
- Chadhurvedi–Graham (2022)

Riemann invariants

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$$\underline{L}\mathcal{R}_- = 0, \quad L\mathcal{R}_+ = 0$$

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- $\mathcal{R}_\pm = v^1 \pm F(\varrho)$ are **Riemann invariants**
- F determined by equation of state $p = p(\varrho, s)$
- $L = \partial_t + (v^1 + c)\partial_1$
- $\underline{L} = \partial_t + (v^1 - c)\partial_1$
- $c = \sqrt{\frac{\partial p}{\partial \varrho}} = \text{speed of sound} > 0$

Shocks for 1D isentropic compressible Euler

Simple (with $\mathcal{R}_- \equiv 0$) isentropic ($s \equiv 0$) plane waves form shocks through the same **Riccati**-type mechanism as in Burgers' equation $\partial_t \Psi + \Psi \partial_x \Psi = 0$,

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“**transversal convexity**.”
- Picture is qualitatively different compared to Burgers' equation: **Cauchy horizons**.
- Cauchy horizons can rescue uniqueness of classical solutions. So far, this is understood only locally in the regime with transversal convexity.

Maximal globally hyperbolic development for 1D isentropic compressible Euler solutions

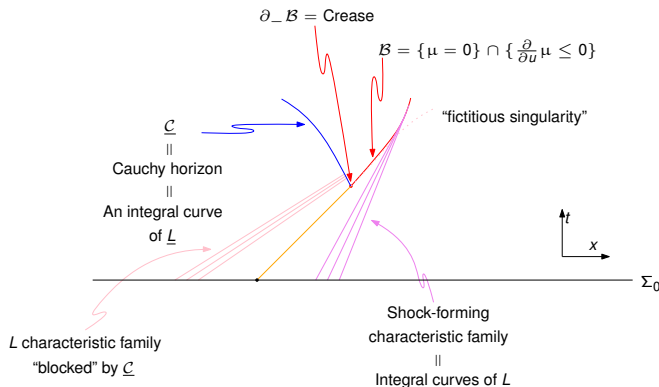


Figure: Local structure of MGHD for \mathcal{R}_+ -dominated 1D isentropic compressible Euler solutions

Multi-dimensions?

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- Geometry plays a key role
- Relies on **energy estimates**, which are very difficult near singularities

Multi- D shocks and singularities

- Majda (1980s)
- Alinhac (late 1990s)
- Christodoulou (2007, 2019)
- Christodoulou–Miao (2014)
- Miao–Yu (2016)
- Holzegel–Luk–Speck–Wong (2016)
- Luk–Speck (2016, 2020s)
- Merle–Raphael–Rodnianski–Szeftel (implosion singularities; 2020s)
- Buckmaster–Cao–Labora–Gómez (more implosions; 2020s)
- Cao–Labora–Gómez–Serrano–Shi–Staffilani (non-radial implosions; 2023)
- Abbrescia–Speck (2020s)
- Buckmaster–Iyer (2020s)
- Buckmaster–Drivas–Shkoller–Vicol (2020s)
- Ginsburg–Rodnianski (pre-print)
- (Luo–Yu) (irrotational rarefaction waves in $2D$)
- Anderson–Luk (pre-print on Einstein–Euler)
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With Abbrescia, for **open sets** of data in $3D$, we have given the first complete description of the **structure of the singular set, including a connected component of the ‘first singularity’, and the Cauchy horizon**

Infinite density of the characteristics

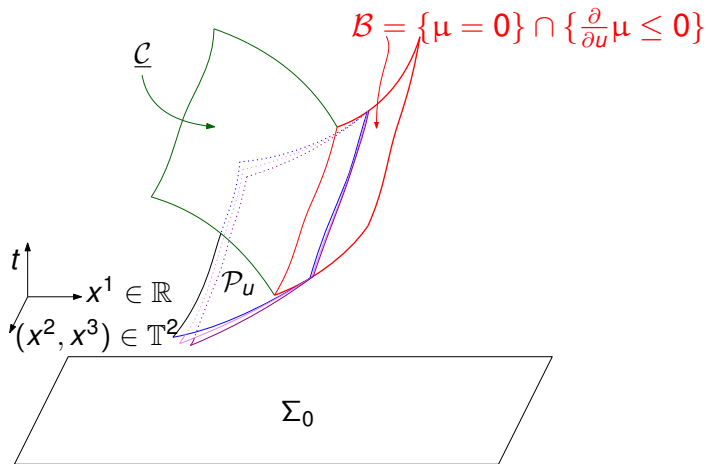


Figure: Infinite density of the characteristics \mathcal{P}_u on \mathcal{B}

New results with L. Abbrescia

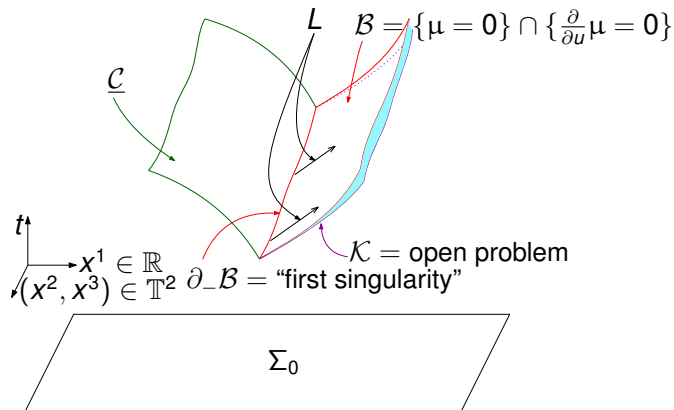


Figure: A localized subset of the maximal classical development and the shock hypersurface in Cartesian space

Acoustical metric

The acoustical metric is tied to sound wave propagation.

Definition (The acoustical metric and its inverse)

$$\mathbf{g} := -dt \otimes dt + c^{-2} \sum_{a=1}^3 (dx^a - v^a dt) \otimes (dx^a - v^a dt),$$

$$\mathbf{g}^{-1} := -\mathbf{B} \otimes \mathbf{B} + c^2 \sum_{a=1}^3 \partial_a \otimes \partial_a$$

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Material derivative vectorfield \mathbf{B} is \mathbf{g} -timelike and thus **transverse** to acoustically null hypersurfaces:

$$\mathbf{g}(\mathbf{B}, \mathbf{B}) = -1$$

Acoustic eikonal function

Definition (The acoustic eikonal function)

The acoustic eikonal function u solves:

$$\begin{aligned} (\mathbf{g}^{-1})^{\alpha\beta} \partial_\alpha u \partial_\beta u &= 0, \\ u|_{t=0} &= -x^1, & \partial_t u &> 0 \end{aligned}$$

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Definition (Geometric coordinates)

We refer to (t, u, x^2, x^3) as the **geometric coordinates**.

Inverse foliation density

Definition

$$\mu := -\frac{1}{(\mathbf{g}^{-1})^{\alpha\beta}\partial_\alpha t\partial_\beta u} = \frac{1}{\mathbf{B}u}$$

Can show:

$$\mu|_{t=0} \approx 1$$

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$\mu = 0$ signifies a shock (infinite density of characteristics and blowup of ∂u)

Proof philosophy

Big idea (Alinhac and Christodoulou): Solution remains rather smooth in geometric coordinates

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$$\bullet \partial_\alpha \sim \frac{1}{\mu} \frac{\partial}{\partial u} + \frac{\partial}{\partial t} + \frac{\partial}{\partial x^2} + \frac{\partial}{\partial x^3}$$

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- $\partial_\alpha \sim \frac{1}{\mu} \frac{\partial}{\partial u} + \frac{\partial}{\partial t} + \frac{\partial}{\partial x^2} + \frac{\partial}{\partial x^3}$
- Hence, $\mu = 0$ represents a degeneracy between Cartesian and geometric partial derivatives

Infinite density of the characteristics

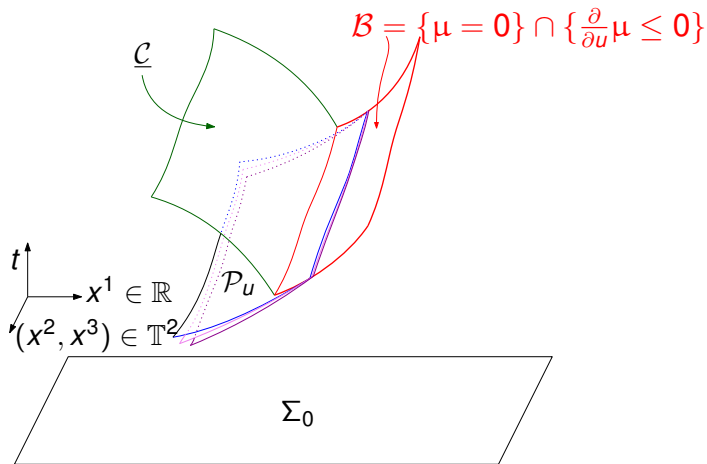


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Strictly convex sub-regime

The strictly convex sub-regime is easier to study:

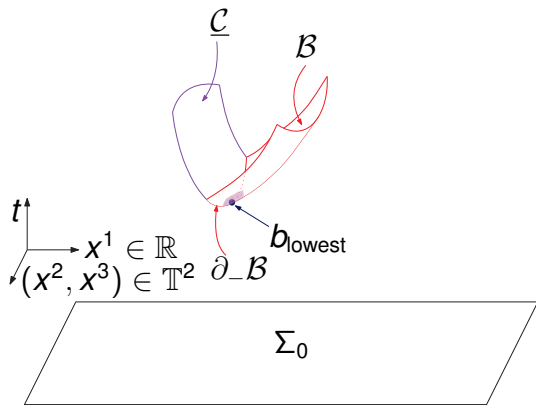


Figure: Strictly convex crease, singular boundary, and Cauchy horizon in Cartesian coordinate space

Null vectorfields

Definition

Null vectorfields

$$\begin{aligned} L_{(geo)}^\alpha &:= -(\mathbf{g}^{-1})^{\alpha\beta} \partial_\beta u, \\ L^\alpha &:= \mu L_{(geo)}^\alpha \end{aligned}$$

Easy to see:

$$Lt = 1$$

In plane symmetry, L agrees with the vectorfield defined explicitly in terms of Riemann invariants

Vectorfield frames constructed from u

Definition (Frame vectorfields)

- X is Σ_t -tangent, left-pointing, satisfies $\mathbf{g}(X, X) = 1$, and \mathbf{g} -orthogonal to $\ell_{t,u} := \Sigma_t \cap \mathcal{P}_u$
- $\check{X} := \mu X$ (satisfies $\check{X}u = 1$)
- For $A = 2, 3$, $Y_{(A)} := \mathbf{g}$ -orthogonal projection of (rectangular) ∂_A onto $\ell_{t,u}$

Definition (Frame adapted to the characteristics)

The **rescaled frame** is:

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Big ideas:

- Derive **regular** estimates relative to the rescaled frame
- Shows that the solution and its $\{L, \check{X}, Y_{(2)}, Y_{(3)}\}$ -derivatives remain **rather smooth** (equivalently, smooth w.r.t. (t, u, x^2, x^3) and smooth in directions tangent to \mathcal{P}_u)

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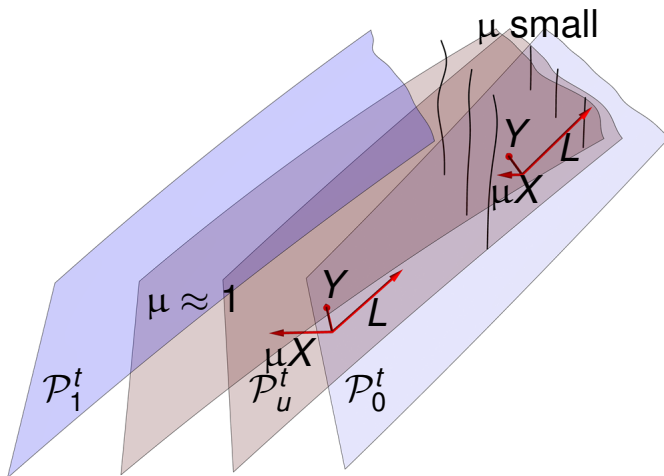
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- Big technical difficulty: **High order geometric energies can blow up** as $\mu \downarrow 0$: $\mathbb{E}_{\text{Top}} \lesssim \mu^{-10}$, $\mathbb{E}_{\text{Top-1}} \lesssim \mu^{-8}$, \dots , $\mathbb{E}_{\text{Mid}} \lesssim 1$

A picture of the dynamics



Statement of main results

Theorem (JS and L. Abbrescia)

*Fix a 1D simple, isentropic shock-forming background solution satisfying the **transversal convexity** condition*

$$\frac{\partial^2}{\partial u^2} \mu|_{\{\mu=0\}} > 0.$$

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- Writing-in-progress: we give a complete description of a neighborhood of the **Cauchy horizon** $\underline{\mathcal{C}}$ that includes the entire crease.
- In total, we reveal a portion of the **maximal (classical) globally hyperbolic development**, including a neighborhood of the boundary.

New results with L. Abbrescia

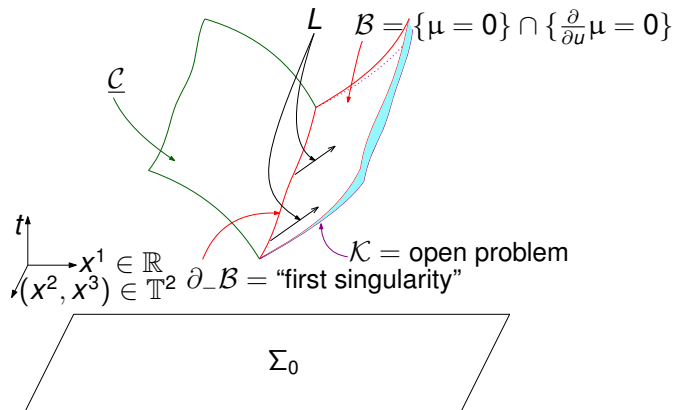
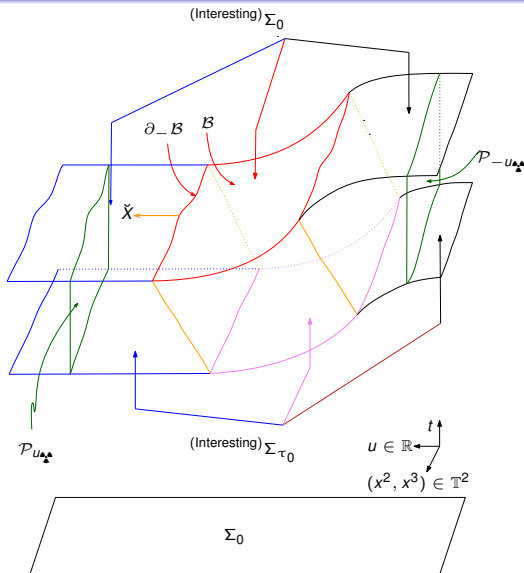


Figure: A localized subset of the maximal classical development and the shock hypersurface in Cartesian space

The crease and the singular boundary



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Then \underline{u} is propagated via the eikonal equation:

$$(\mathbf{g}^{-1})^{\alpha\beta} \partial_{\alpha} \underline{u} \partial_{\beta} \underline{u} = 0$$

The Cauchy horizon region

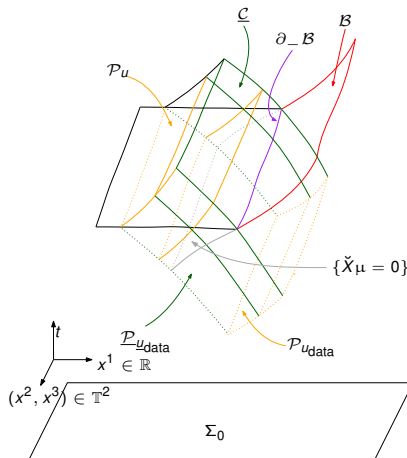


Figure: The Cauchy horizon region in Cartesian space

Connection to wave equations

In isentropic plane symmetry, the equations reduce to $L\mathcal{R}_{(+)} = 0$, $\underline{L}\mathcal{R}_{(-)} = 0$. In particular:

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- There are many tools for geometric wave equations
- Also useful for low-regularity well-posedness

New formulation of 3D compressible Euler

Theorem (J. Luk–JS; M. Disconzi–JS in relativistic case)

Consider smooth compressible Euler solutions in 3D. For $\Psi \in \vec{\Psi} := (\varrho, v^1, v^2, v^3, s)$, we have, *schematically*:

$$\begin{aligned} \square_{\mathbf{g}(\vec{\Psi})} \Psi &= \nabla \times \left(\frac{\nabla \times \mathbf{v}}{\varrho} \right) + \operatorname{div} \nabla s \\ &\quad + \mathbf{g}\text{-null forms}, \\ \mathbf{B} \left(\frac{\nabla \times \mathbf{v}}{\varrho} \right) &= \nabla \vec{\Psi} \cdot \left(\frac{\nabla \times \mathbf{v}}{\varrho} \right) + \nabla \vec{\Psi} \cdot \nabla s, \\ \mathbf{B} \nabla s &= \nabla \vec{\Psi} \cdot \nabla s \end{aligned}$$

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→ With L. Abbrescia, we derived suitable “elliptic-hyperbolic” identities for $\frac{\nabla \times \mathbf{v}}{\varrho}$ and ∇s on **arbitrary globally hyperbolic domains** for 3D compressible Euler solutions

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Would require the development of **new geometry**.